UCLA Department of Electrical Engineering EE170C – Photonic Sensors and Solar Cells Spring 2015 Midterm, April 29 2015, (1:40 minutes)

Name	Student number	*
This is a closed book exam – yo	ou are allowed 1 page of notes (front+back).	

Check to make sure your test booklet has all of its pages – both when you receive it and when you turn it in.

Remember – there are several questions, with varying levels of difficulty, be careful not to spend too much time on any one question to the exclusion of all others.

Exam grading: When grading, we focusing on evaluating your level of understanding, based on what you have written out for each problem. For that reason, you should make your work clear, and provide any necessary explanation. In many cases, a correct numerical answer with no explanation will not receive full credit, and a clearly explained solution with an incorrect numerical answer will receive close to full credit. CIRCLE YOUR FINAL ANSWER.

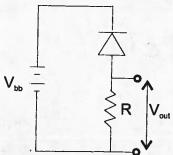
If an answer to a question depends on a result from a previous section that you are unsure of, be sure to write out as much of the solution as you can using symbols before plugging in any numbers, that way at you will still receive the majority of credit for the problem, even if your previous answer was numerically incorrect.

Please be neat – we cannot grade what we cannot decipher.

	Topic	Max Points	Your points
Problem 1	Photodetection and APD	30	Tour points
Problem 2	Detection limits	40	
Problem 3	Photodiodes	30	
Total		100	

1. Detection (30 points)

Consider a Si photodiode with the following characteristics: area= 0.2 mm², quantum efficiency η =0.7 (measured at a wavelength of λ =900 nm), and dark current I_d =1 pA. The stray capacitance is 1 pF...



(a) (10 points) Find the minimum detectable power at 900 nm assuming that your detection bandwidth is B=1Hz.

Find NEP assuring R doesn't contribute Johnson noise (quantum limited regime).

1 = ONEP NEP JZeIB

For $\lambda = 900 \text{ nm}$ & by = 1.38 V $\sigma = \frac{9e}{hv} = 0.50 \text{ A/w}$ $NEP = 1.13 \times 10^{-15} \text{ W}$

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(b) (10 points) What value should you choose for the load resistance R so that you are in the quantum limited detection regime?

Enor in wording of question:

Should have asked for clark current limited detection regime. Quantum limited detection implies date when it is negligible. Feel credit for alternate answer:

MAM

in = ZeB(Iph+Id) + YKIB

(c) (10 points) Now consider that this photodiode is replaced by an APD with the same characteristics, except with gain M=100 and noise factor F=1.3. (You may assume that the dark current is also multiplied by the same factor). How do your answers to (a) and (b) change? Given your answer, what is the primary advantage of using an APD for detection of low –level light.

How does a) change? Id-> ROW MID $I = \frac{O M NEP}{\sqrt{2e I_a M^2 BF}}$ $NEP = \sqrt{2e I_a F-B}$ $= 1.28 \times 10^{-15} \text{W}$

The Mouters cancel, so NEP is only Slightly increased by a factor of JF

b)? 4kTB << 2e Tam BF

R>>> 2kT

M2e Ta F

R>>> 4.0x106_12

The amplification drastically reduces the value as of R needed to reach appears that the detection limited by dark current (factor of 1/m²),

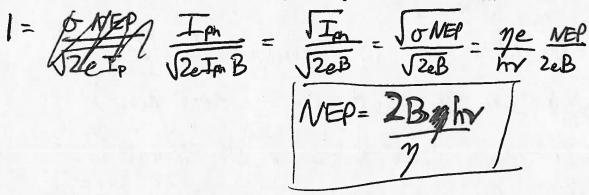
This will perhase increase detection bandwidth by factor of M² from & 31+2 to LC 40 KHz

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2. (40 points) Detection limits

(a) (15 points) In class, we usually assumed that the minimum detectable power (i.e. NEP) was either limited by background signal current or dark current. Consider the case if the detector was cooled and such that I_d is negligible, and there is no significant background current. Derive an expression for the noise equivalent power NEP in this case.



(b) (10 points) Consider a photodetector for which $\eta=1$, and for light with $hv=1\,\text{eV}$. What is the value of NEP in this case (assume $B=1\,\text{Hz}$). How small must I_d be to reach this limit?

(c) (15 points) Explain what the BLIP Detectivity figure of merit is. Is BLIP detectivity more important for photodetectors at $\lambda=1$ µm or $\lambda=10$ µm? Why?

Backgrand Limited Indirect Performance: This is the

Value of NEP or D* produced by the

background sign photocurrent (+ its shot noise)

When observing a back scene & its thermal radiation. It is more important at 10 um because the blackbody radiation is much stronger (at 300K) at 10 um than lum

- 3. (30 points) Photodiodes
 - (a) Why are III-V photodiodes typically faster (larger cutoff frequency) than silicon based structures? Give two reasons.
 - -> Larger absorption coeff of allows thinner absorption regions + shorter transit times
 - -> III-I have larger mobility of Saturation velocities.

(b) pin photodiodes often have a dark current value that is less dependent on bias voltage than p-n photodiodes. Explain why this might be the case.

The lots photodiode dark current is primarily determined by the thermal generation Id=egiA(L+Ln+w) when wis the depletion region width.

For apm diple wolfVI in reverse bias, which leads to an increase in dark ament Is a JIVI with reverse bias.

For a pin, most of the voltage drop is across the "i" layer, which is independent of bias. This leads to a dependence of w that is slower than JUI.

(c) Explain the tradeoff between high speed operation and quantum efficiency in a conventional photodiode.

The ultimate limit for high speed is the trunsit time limit: fed - ITT-Zd where Ta = Warg

Where wis the depletion thickness. However, in III-II's nearly all of the absorption takes place in the clepthon region. If you decrease W, you also decrease the quantum efficiency as the length available for absorption decreases.

n=1-e-aw adw. (dwell

This leads to an approximate constant bardwith left product for surface incident photodiodes.

frage The dw = Vand.

Blackbody radiation energy density in cavity:

$$\rho(v) = \frac{8\pi h v^3}{c^3} \frac{1}{e^{hv/k_B T} - 1}$$

$$(J Hz^{-1} m^{-3})$$

Spectral Radiance from Blackbody surface:

$$R(v) = \frac{2hv^3}{c^2} \frac{\cos \theta}{e^{hv/k_BT} - 1}$$
 (W ster-1 Hz-1 m-2)

$$R(v) = \frac{2hv^3}{c^2} \frac{\cos \theta}{e^{hv \cdot k_B T} - 1} \qquad \text{(W ster^-1 Hz^{-1} m^{-2})} \qquad \qquad R(\lambda) = \frac{2hc^2}{\lambda^5} \frac{\cos \theta}{e^{hc/\lambda k_B T} - 1} \qquad \text{(W ster^-1 m^{-2})}$$

Radiant Emittance from Blackbody surface: $E = \frac{\pi^2 (k_B T)^4}{60 \, c^2 t^3} = \sigma_B T^4$ (W m⁻²)

Semiconductors

Fermi-Dirac distribution for electrons:

$$f(E) = \frac{1}{e^{(E-E_F)/kT} + 1}$$
, for holes: $f(E) = 1 - \frac{1}{e^{(E-E_F)/kT} + 1}$,

Equilibrium Carrier concentrations

$$n_0 = \int_{E_{-}}^{\infty} f(E)N(E)dE$$

Equilibrium Carrier concentrations in non-degenerate limit $(E_C - E_F \gg kT)$ and $E_F - E_V \gg kT$.

$$n_0 = N_C e^{-(E_C - E_F)/kT}$$
 , $p_0 = N_V e^{-(E_F - E_V)/kT}$

$$N_C = 2\left(\frac{2\pi m_n^* kT}{h^2}\right)^{3/2}$$
, $N_V = 2\left(\frac{2\pi m_p^* kT}{h^2}\right)^{3/2}$

$$n_i = \sqrt{n_0 p_0} = \sqrt{N_C N_V} e^{-E_R/2kT}$$

Intrinsic Fermi Level

$$E_i = \frac{kT}{2} \ln \left(\frac{N_v}{N_C} \right) + \frac{E_v + E_C}{2}$$

Conductivity of semiconductor:

$$\sigma = ne\mu_n + pe\mu_p$$

Semiconductor electron/hole mobility

$$\mu_n = \frac{e\overline{t_n}}{m_n^*} \qquad \qquad \mu_p = \frac{e\overline{t_p}}{m_n^*}$$

Einstein relation for diffusion coeff: $D = \frac{kT}{\rho}\mu$

Diffusion length $L = \sqrt{D\tau}$

Ideal p-n diode (assumes abrupt junction)

Contact potential

$$V_0 = \frac{kT}{e} \ln \left(\frac{N_A N_D}{n_i^2} \right),$$

Depletion (transition) region width

$$W = \sqrt{\frac{2\epsilon \left(V_0 - V\right)}{e} \left(\frac{1}{N_A} + \frac{1}{N_D}\right)}, \qquad x_{n0} = W \frac{N_A}{N_A + N_D}$$

Maximum Electric field within depletion region

$$\mathcal{E}_{\text{max}} = \sqrt{\frac{2e(V_0 - V)}{\epsilon} \left(\frac{N_A N_D}{N_A + N_D} \right)}$$

Ideal diode I-V:

$$I = eA \left(\frac{D_p}{L_p} p_n + \frac{D_n}{L_n} n_p \right) \left(e^{eV/k_BT} - 1 \right) = I_0 \left(e^{eV/k_BT} - 1 \right)$$

Depletion (junction) Capacitance

$$C_{j} = \epsilon A \sqrt{\frac{e}{2\epsilon (V_{0} - V)} \frac{N_{A} N_{D}}{N_{A} + N_{D}}} = \frac{\epsilon A}{W}$$

Transit time (average):

$$\tau_d \approx \frac{w}{2} \left(\frac{1}{v_e} + \frac{1}{v_h} \right)$$

Noise and photodetection

$$V(\omega) = \int_{-\infty}^{\infty} v(t)e^{-j\alpha t}dt$$

Fourier Transform:

$$v(t) = \frac{1}{2\pi} \int_{-\infty}^{\infty} V(\omega) e^{j\omega t} d\omega$$

Single sided spectral density over interval T:

$$S_T(v) = \frac{2}{T} \big| I_T(\omega) \big|^2$$

Johnson noise

$$\langle i_J^2 \rangle = 4k_B TB/R$$
 Shot noise

$$\langle i_s^2 \rangle = 2eIB$$

Shot noise with avalanche gain M

$$\langle i_s^2 \rangle = 2eIBM^2F$$

Signal to noise ratio:

$$SNR_{P} = \frac{I_{ph}^{2}}{\langle i_{n}^{2} \rangle}$$

$$SNR_{P} = \frac{I_{ph}^{2}}{\left\langle i_{n}^{2} \right\rangle}$$
 $SNR_{I} = \frac{I_{ph}}{\sqrt{\left\langle i_{n}^{2} \right\rangle}}$

Detectivity

$$D^* = \frac{\sqrt{AB}}{NEP}$$

Constants (SI units):

$$\varepsilon_0 = 8.85 \times 10^{-12} \text{ F/m} \text{ (or C}^2 \text{ N}^{-1} \text{ m}^{-2}\text{)}$$

 $e = 1.6 \times 10^{-19} \text{ C}$

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$$h = 6.626 \times 10^{-34} \text{ J s}$$

$$\mu_0$$
=4 π x10⁻⁷ H/m (or N A⁻²)
 k_B = 1.38e⁻²³ J/K
 \hbar = 1.055x10⁻³⁴ J s

$$k_B = 1.38e^{-23} \text{ J/K}$$

$$\hbar = 1.055 \times 10^{-34} \text{ J s}$$

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